

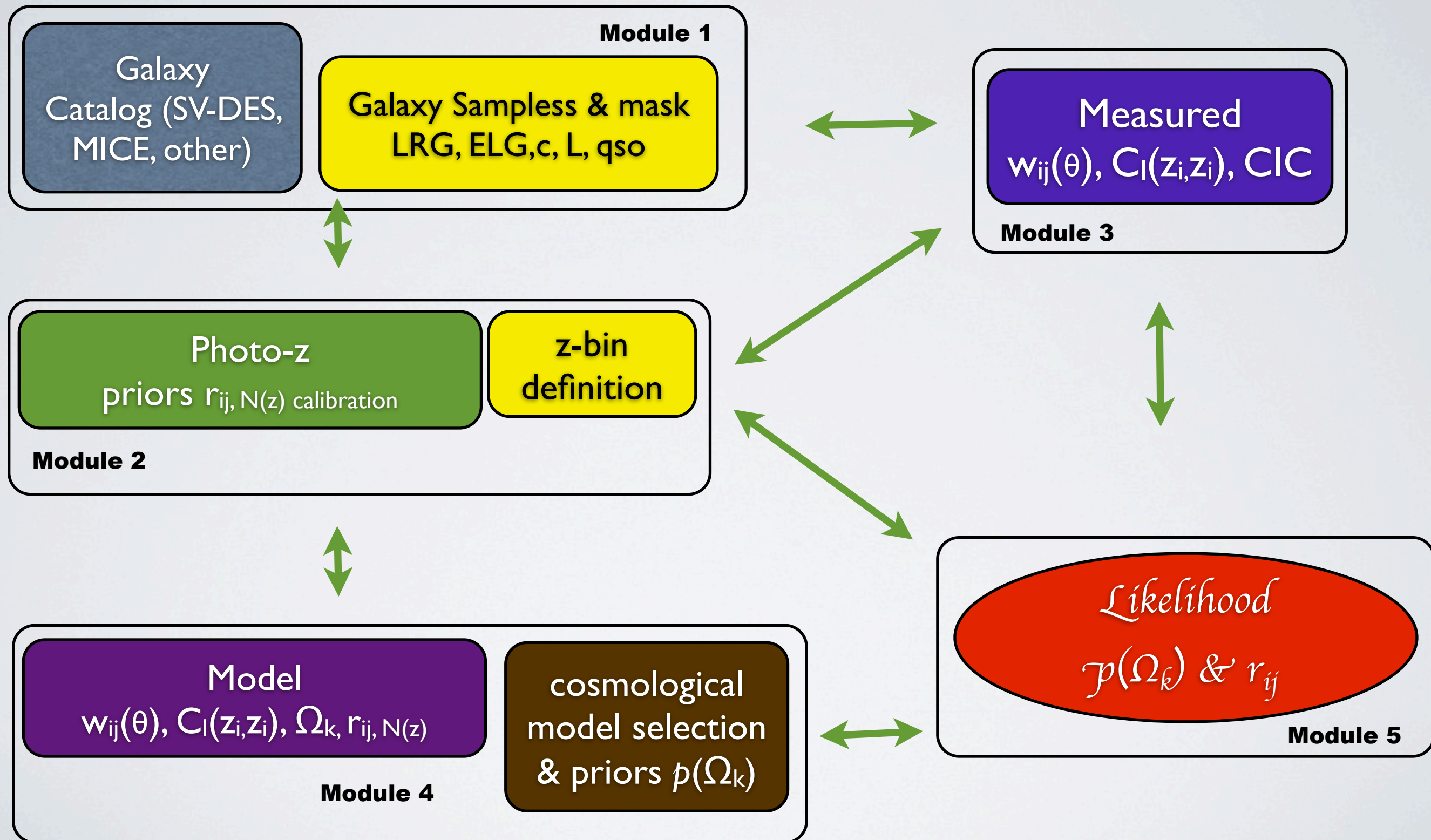
# Early Galaxy Clustering Measurements in SV data

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# An end-to-end pipeline for cosmological analysis

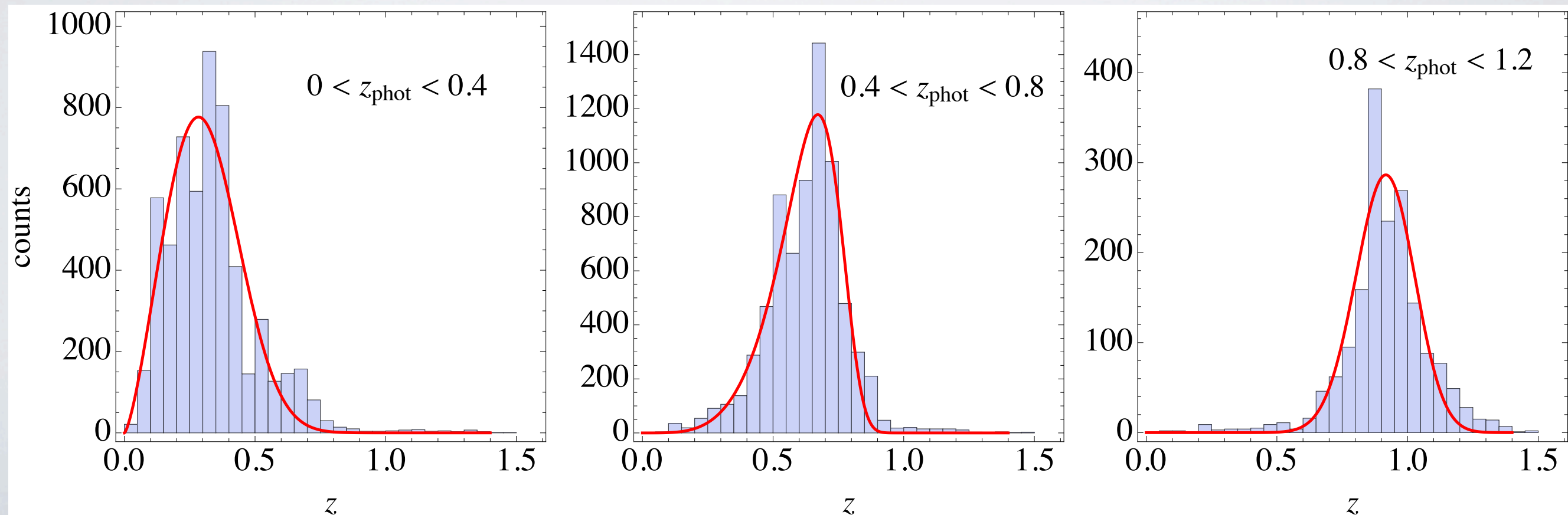


# Clustering analysis in the 4 calibration fields

- VVDS Deep 02hr
- CDFS
- VVDS Deep 14hr
- Cosmos

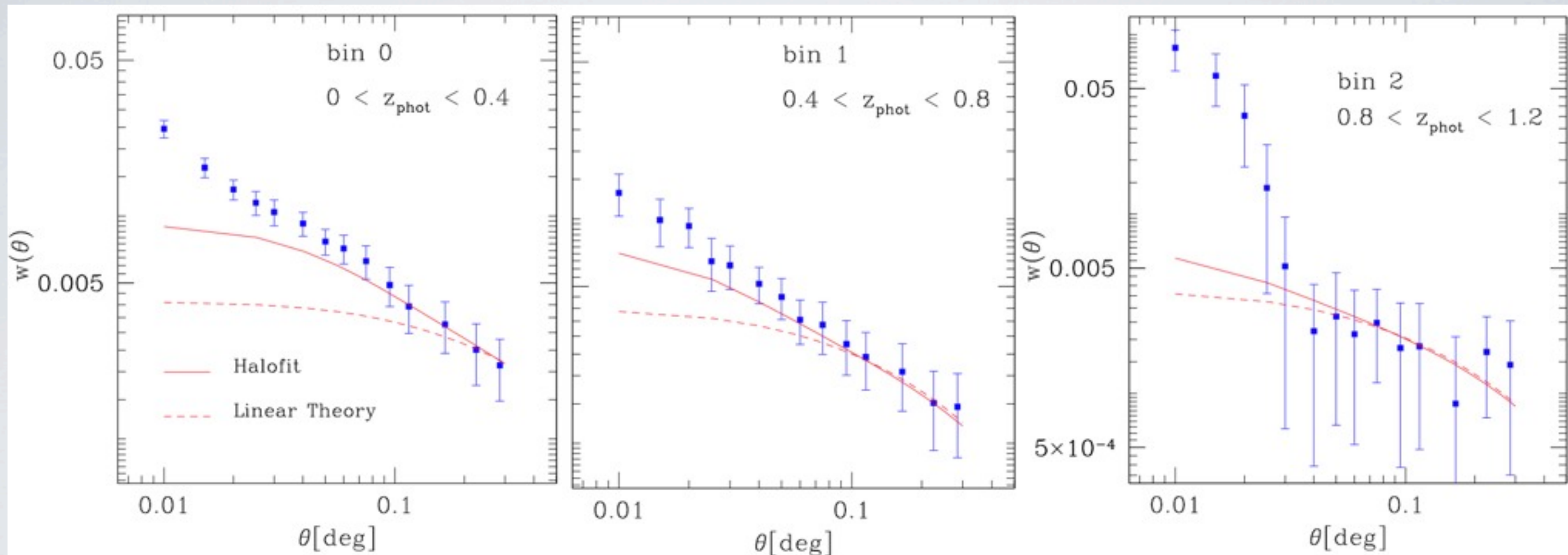
- Roughly 1 sq deg each field and 500,000 objects
- Mag limit selection ( $18 < i\_mag\_auto < 24$ ) and star-gal separation ( $spread\_model > 0.002$ )
- We run 2 different photo-z codes ANNz and BPZ

## BPZ results : redshift distributions in 3 bins



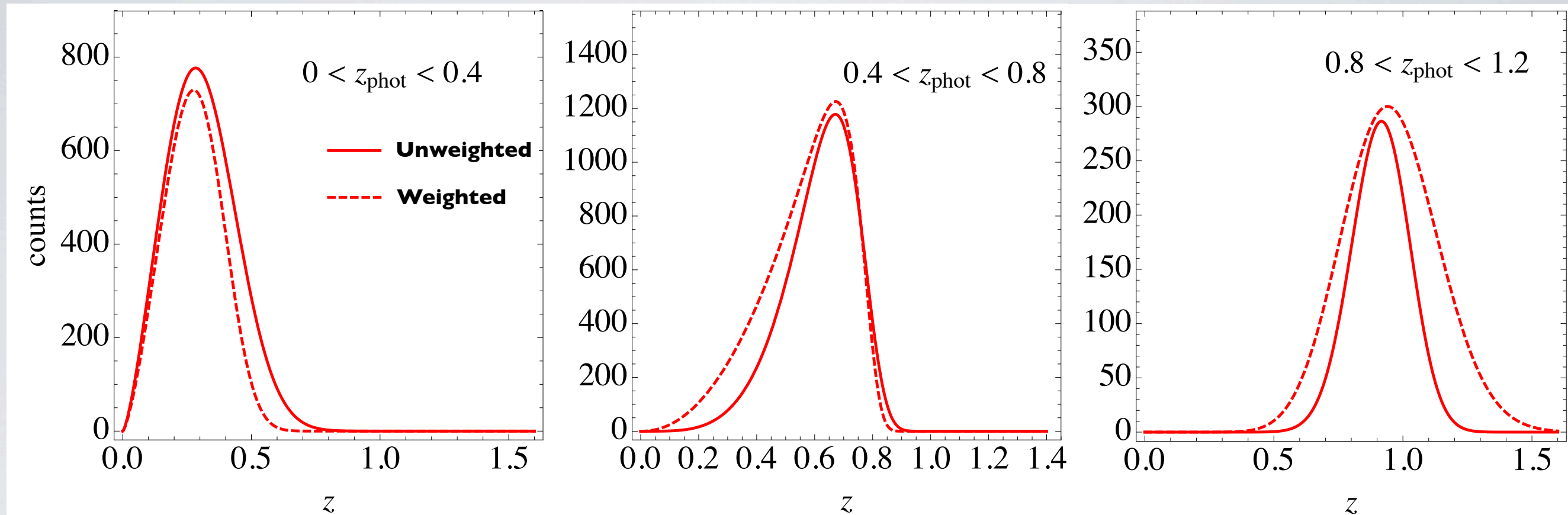


## BPZ results : small scale clustering in 3 bins

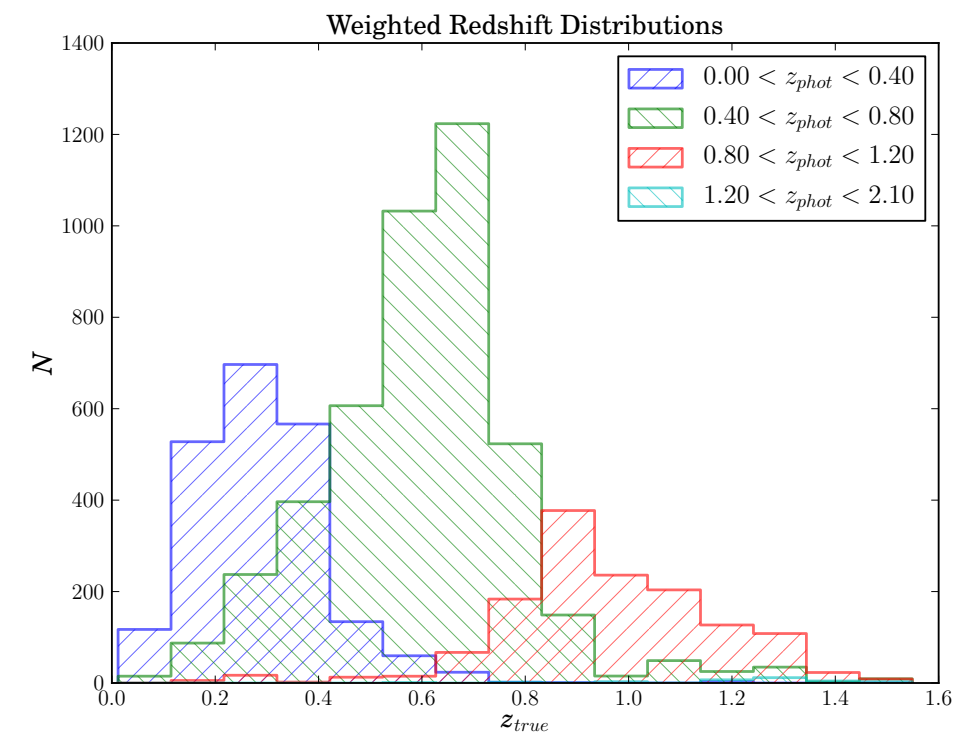


► Nonlinear Effects start to kick in at least at 0.1 deg, the clustering excess over the model towards smaller angles could be due to nonlinear biasing, systematics, ..

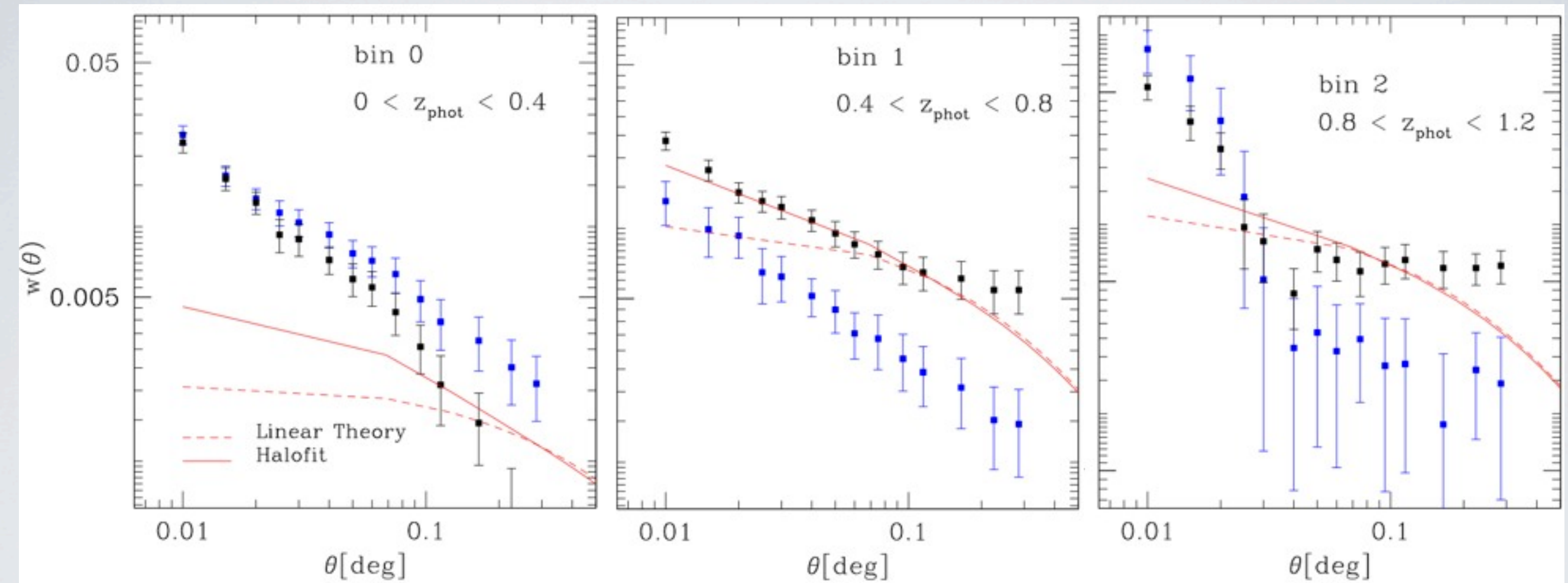
## ANNz : weight the spectroscopic sample to estimate $N(z)$



► Distributions are broader (but more realistic photo- $z$ )



## ANNz & BPZ results : small scale clustering in 3 bins

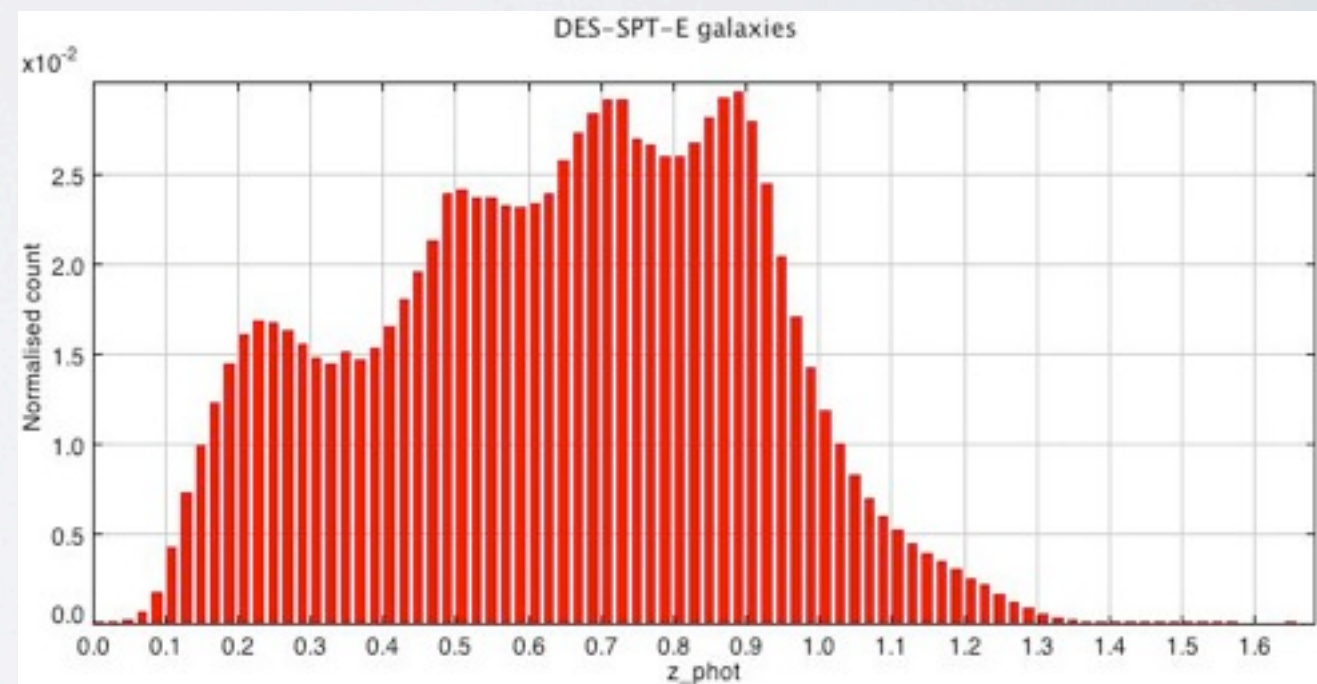
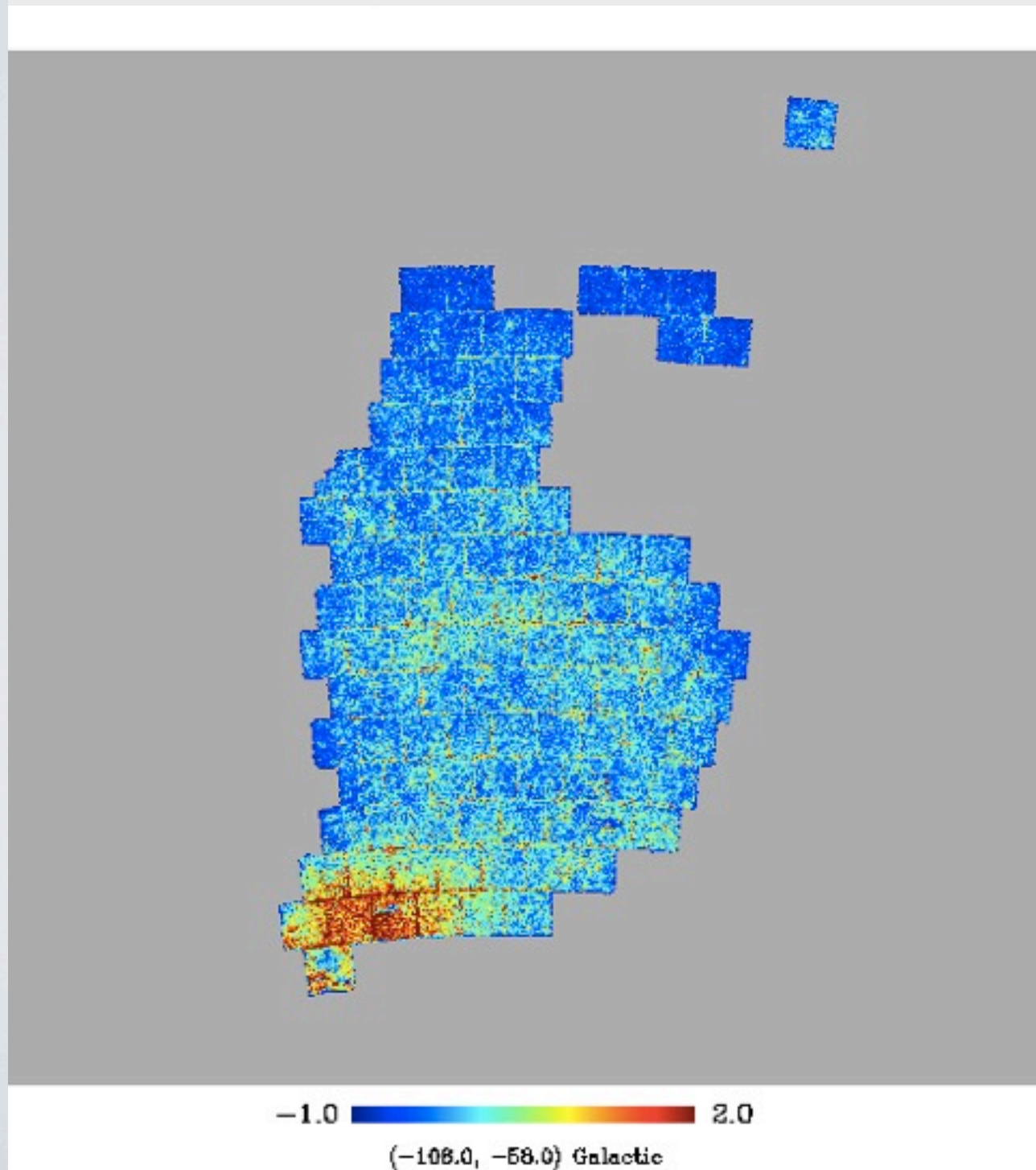


- Black symbols are ANNz , Blue same BPZ (as previous slides)
- A useful comparison; clearly shows problems in the data and/or reduction.



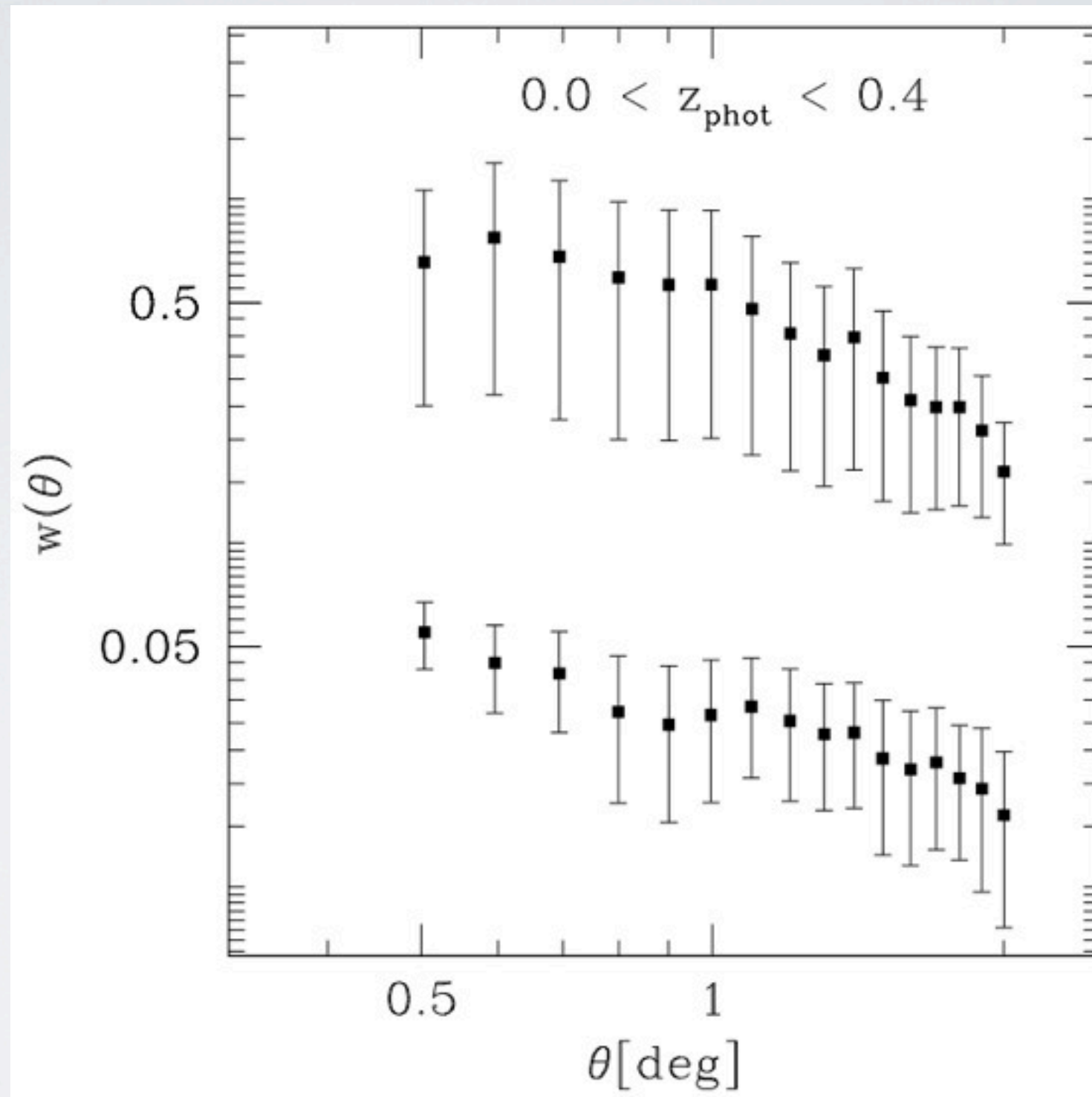
# Clustering analysis in SPT-E field

- Similar selection as before ( $i\_mag\_auto < 24$ , and  $i\_spread > 0.002$ ), 1.4 million objects
- Notice the crowded [source contamination](#) at  $DEC < -60$ . This is probably due to stars from the LMC, a problem for [low-z](#) co-adds



ANNZ results

➤ Impact on Clustering for low-z bin

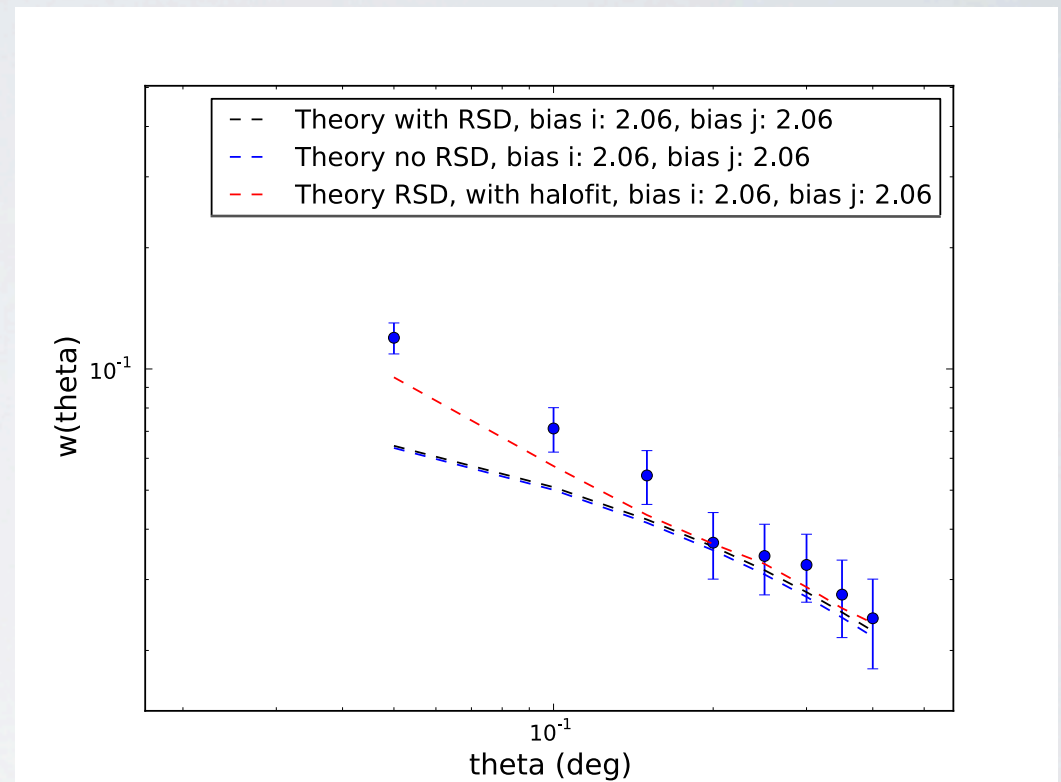
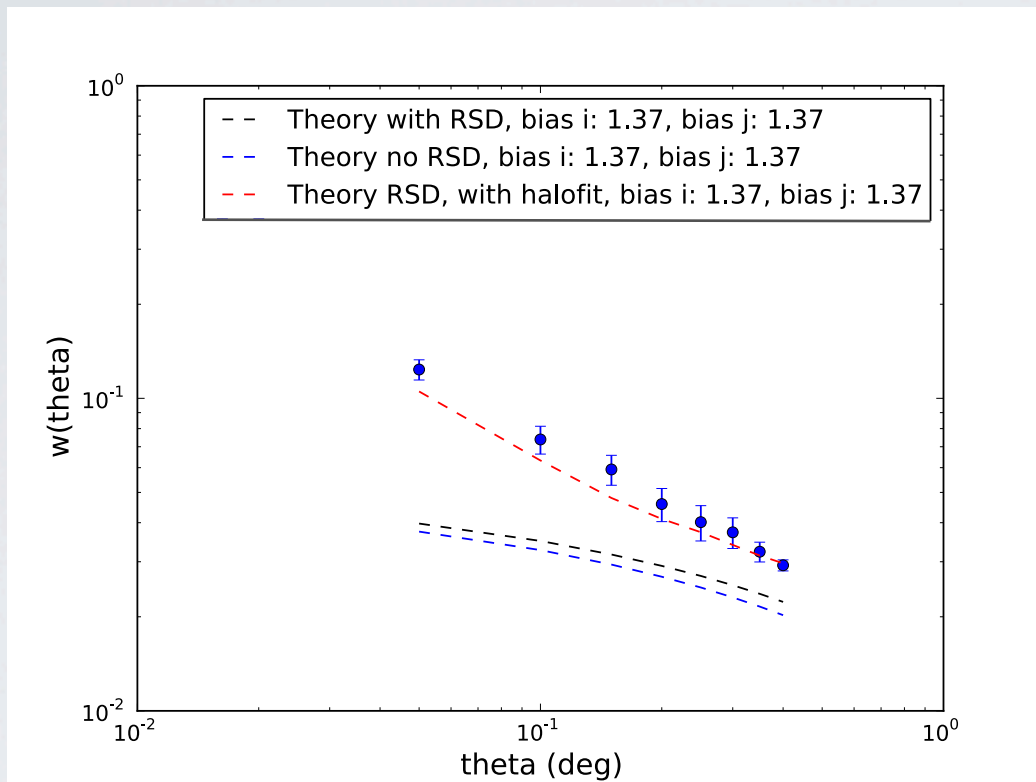




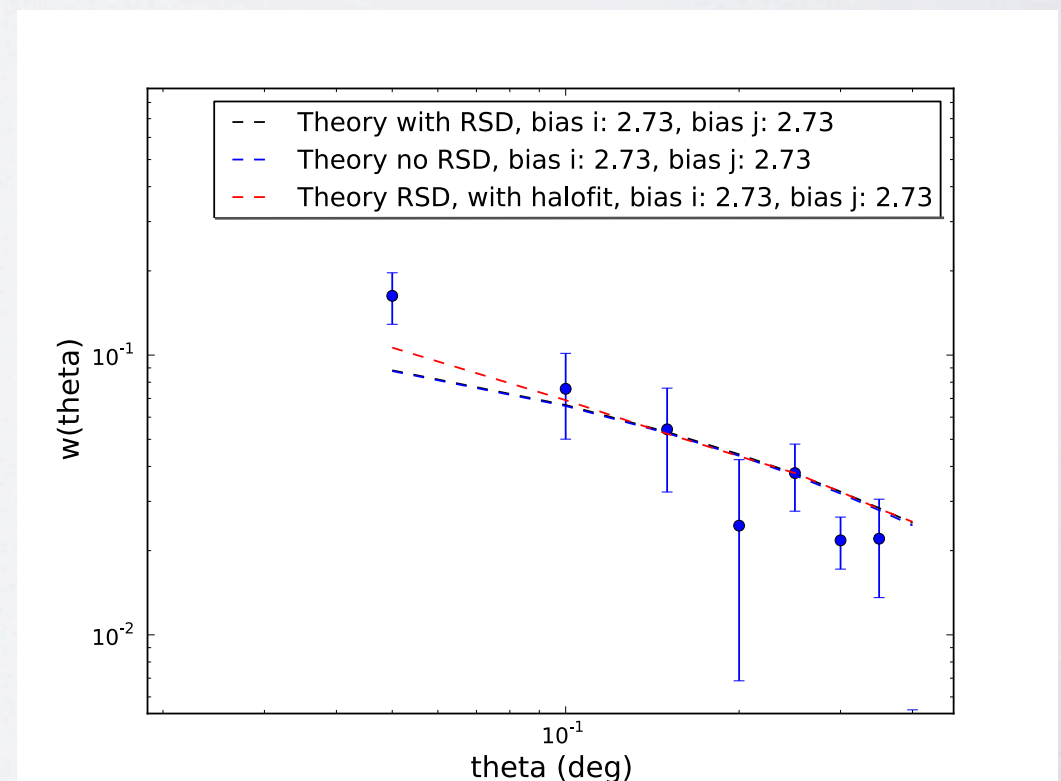
# Clustering analysis in SPT-E field

cutting out the DES < -6l region

Autocorrelations using BPZ



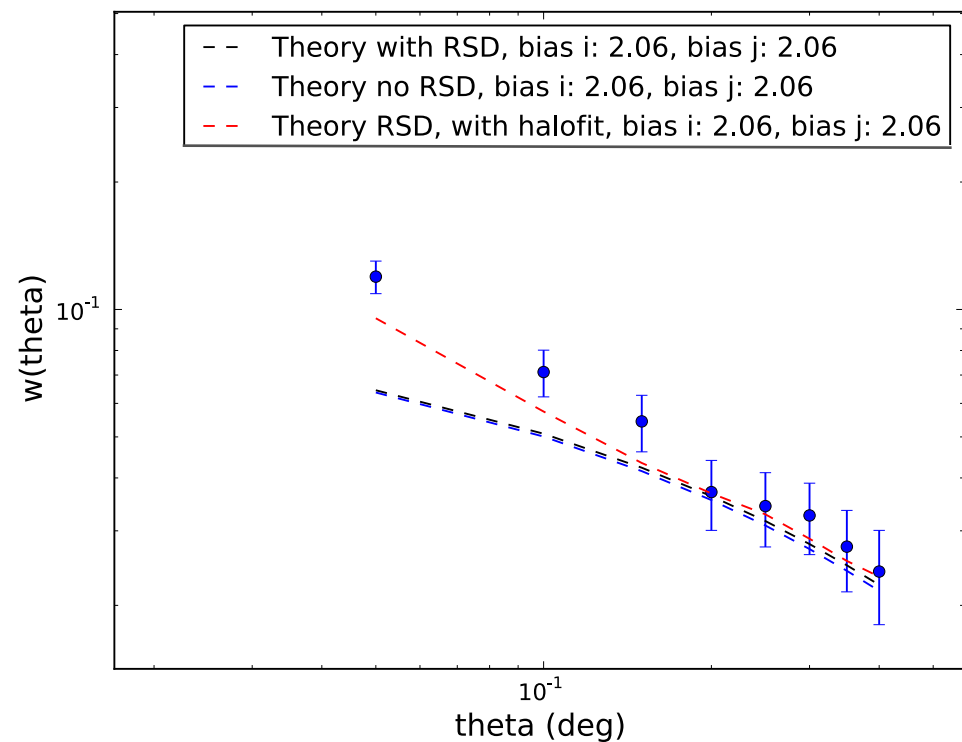
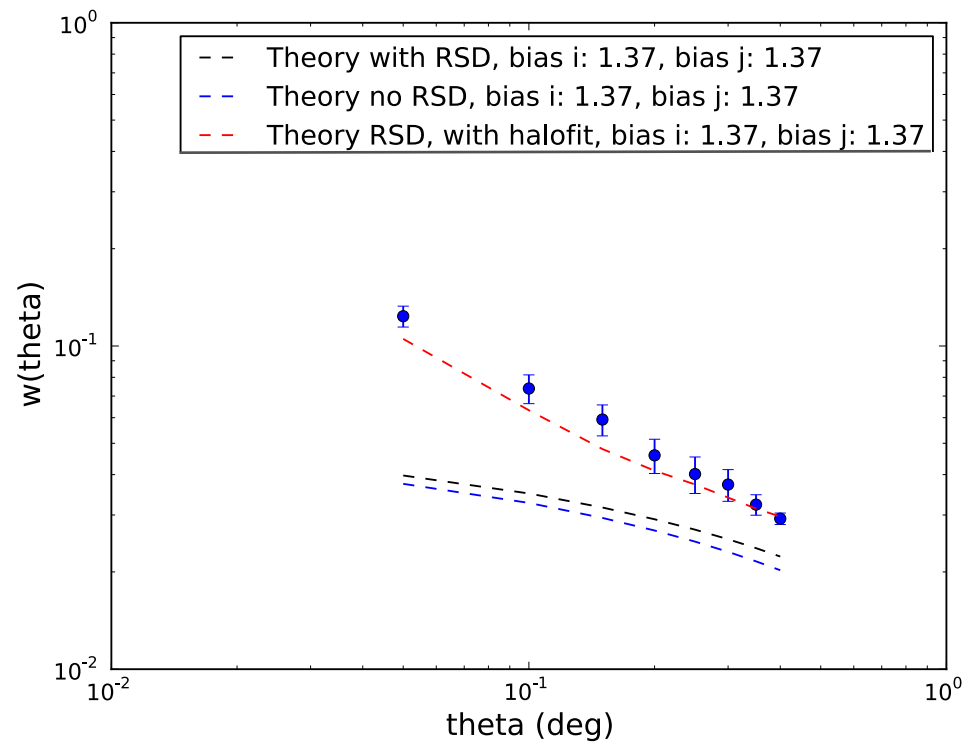
- No training set in the field
- Are the best-fit biases reasonable?



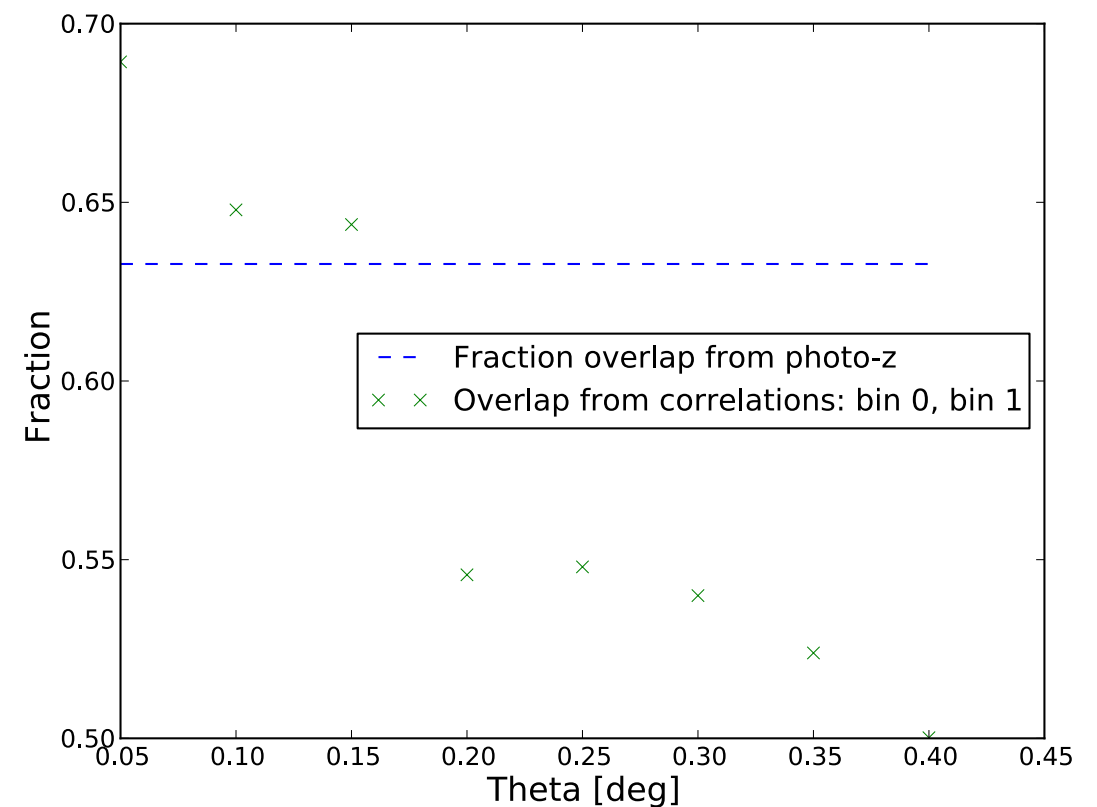
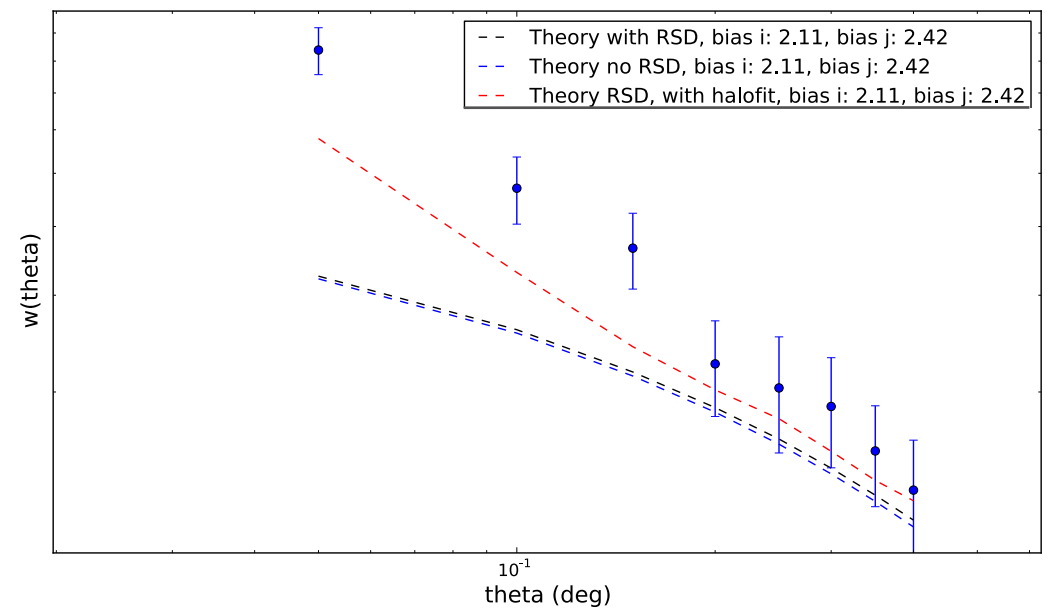
# Cross-correlations

SPT-E region (dec > -61), z 0.0-0.4 vs. 0.4-0.8, BPZ

Autocorrelations bin 0 & 1

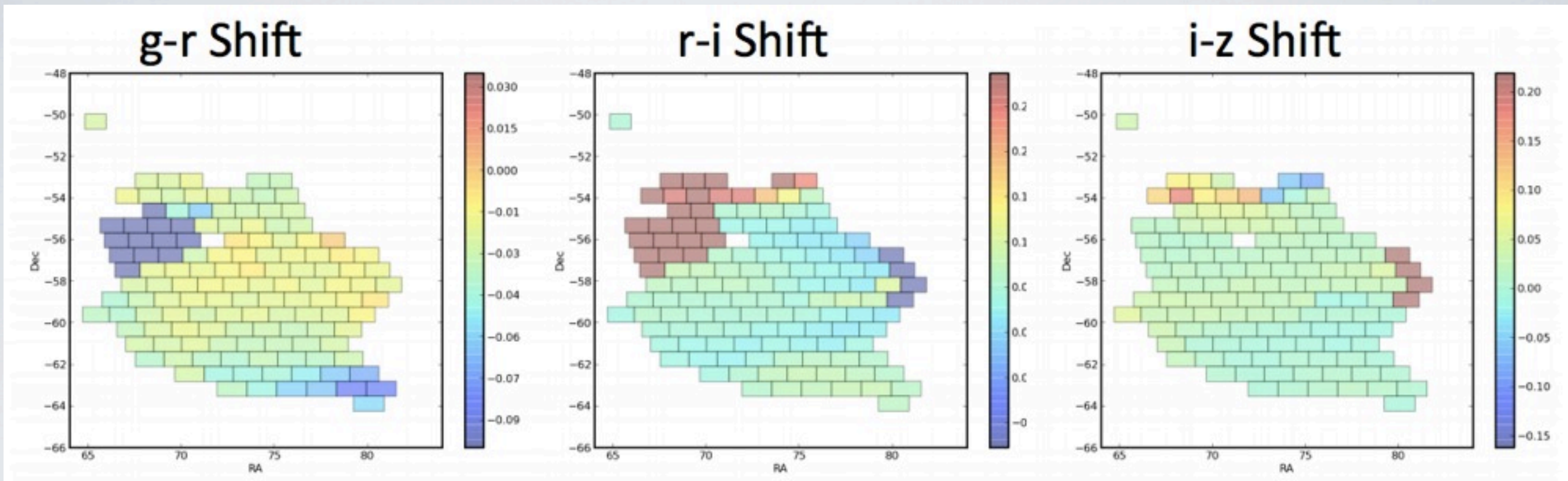


Cross-correlation bin 0 with 1



# Clustering analysis in SPT-E field

effects of calibration: stellar locus measurements



From Bob Armstrong's Talk, Calibration session on Tuesday

➤ Color shifts of a few percent → Trouble in our mag-limited samples!



# Summary

- We have been studying the SV fields
  - Good to compare the 4 calibration fields with the SPT-E
- Developing an analysis pipeline: photo-z, clustering measurement, modeling, likelihood
  - Now at prototype stage
- Data are not like simulations!
  - Need to understand the “features” & make more quality cuts
  - Should give feedback to the DESDM people this spring, before next reduction run.